

Quantum Gravity in Large Extra Dimensions

Emidio Gabrielli
(Helsinki Institute of Physics)

Summary

- **Introduction to Extra Dimensions**
 - Kaluza Klein Theories
 - Unification of Gauge and Gravitational Forces
- **Extra Dimensions and the Hierarchy Problem**
- **Quantum Gravity in Large Extra Dimensions**
- **Experimental Signals at Colliders**
- **Conclusions**

Space Dimensions and Unification

- Connection between unification of forces and number of dim.
- Electric \vec{E} and Magnetic \vec{B} fields unified in continuum 4D space-time (recognized first by Minkowski)

- Four vector potential $A_\mu = (\Phi, \vec{A})$

$$\vec{E} = -\vec{\nabla}\phi - \frac{\partial\vec{A}}{\partial t}, \quad \vec{B} = \vec{\nabla} \times \vec{A}$$

- e.m. fields \vec{E}, \vec{B} embedded in $F_{\mu\nu}$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & B_z & -B_y \\ E_y & -B_z & 0 & B_x \\ E_z & B_y & -B_x & 0 \end{pmatrix}$$

$$L = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu}$$

- char. density and current $\rho, \vec{J} \rightarrow J_\mu = (\rho, \vec{J})$
- Maxwell's eqs $\rightarrow \partial_\mu F^{\mu\nu} = J^\nu$

Unification of E.M. and Gravity

(history of attempts)

- **1912: Gunnar Nordström** proposes gravity with scalar field coupled to trace of energy-momentum tensor T_{μ}^{μ} .
- **1914: G.N.** introduces a 5-dim A_{μ} to describe both EM and Gravity- (work ignored and recently rediscovered).
- **1919: Theodor Kaluza** writes a 5-dim theory for EM and Gravity. Einstein suggests publication 2 years later.
- **1926: Oscar Klein** rediscovers the theory, gives geometrical interpretation and finds **charge quantization**.

Now theory is known as **Kaluza-Klein** theory.

E.M.

Action for free fermions

$$S_{\text{Dirac}} = \int d^4x \bar{\psi} (i\gamma^\mu \partial_\mu - m) \psi$$

invariant under global gauge transformation

$$\psi \rightarrow e^{i\Lambda} \psi$$

Gauge Principle:

E.M. action can be obtained by requiring invariance under **local gauge transformations**
 $\Lambda \rightarrow \Lambda(x)$.

$$S_{\text{EM}} = \int d^4x \left[\bar{\psi} (i\gamma^\mu \mathbf{D}_\mu - m) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right]$$

$$\mathbf{D}_\mu = \partial_\mu - iA_\mu, \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\text{local gauge transf.} = \begin{cases} \psi \rightarrow e^{i\Lambda(x)} \psi \\ A_\mu \rightarrow A_\mu + \partial_\mu \Lambda(x) \end{cases}$$

A_μ is a dynamical field (**photon**)

GRAVITY

In General Relativity the dynamical variable is the space-time metric $g_{\mu\nu}$

- $g_{\mu\nu}$ is a 4×4 symmetric tensor, describing space-time geometry

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

- **Gravitons ($h_{\mu\nu}$):** elementary quantum fluctuation of $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$ around Minkowski (flat) metric $\eta_{\mu\nu} = \text{diag}(1, -1 - 1 - 1)$.
- Dynamics described by **Einstein action**

$$S_G = \frac{1}{16\pi G_N} \int d^4x \sqrt{-g} R(g)$$

- G_N is the Newton's constant
- R is the curvature (function of the metric)
- S_G invariant under general coordinate transformations

Unification of EM and Gravity

Kaluza's idea is to consider General Relativity in **5-dim**

$$\hat{S}_G = \frac{1}{16\pi\hat{G}_N} \int d^5x \sqrt{-\hat{g}} R(\hat{g})$$

where \hat{g}

$$\hat{g}_{MN}(\hat{x}) = \begin{pmatrix} g_{\mu\nu} + k^2\Phi A_\mu A_\nu & k\Phi A_\mu \\ k\Phi A_\mu & \Phi \end{pmatrix}(\hat{x})$$

- dynamical fields $\hat{g}_{MN} \rightarrow g_{\mu\nu}, A_\mu, \Phi$
- \hat{x} is a 5-dim coordinate of space. x_5 is periodic or compactified.

Kaluza-Klein modes

- Assume space is $M_4 \times S_1$

$$x_5 + 2\pi R = x_5$$

- All fields can be expanded in Fourier modes

$$\varphi(x, x_5) = \sum_{n=-\infty}^{\infty} \frac{1}{\sqrt{2\pi R}} \varphi^{(n)}(x) e^{i\frac{nx_5}{R}}$$

- **Equation of motion** (scalar field)

$$\left(\eta_{MN}\partial^M\partial^N + m_0^2\right)\varphi(x, x_5) = 0$$

- **Fourier transform** $\hat{\varphi}^{(n)}$ of each mode (\mathbf{n}) satisfies (in unity $\hbar, c = 1$)

$$\left[E^2 - \vec{\mathbf{p}}^2 - \left(m_0^2 + \frac{\mathbf{n}^2}{R^2}\right)\right]\hat{\varphi}^{(n)} = 0$$

- mass of n^{th} KK mode is

$$m_{\mathbf{n}}^2 = m_0^2 + \frac{\mathbf{n}^2}{R^2}$$

- **Spectrum** = zero mode + KK excitations of mass $m_{\mathbf{n}}^2$

Unification

- Suppose typical energy $E \ll R^{-1} \rightarrow$ only zero modes can be excited.
- Expand the Einstein action in 5-dim, keeping only zero modes and set $\Phi = 1$

$$\hat{S}_G(\hat{g}) = S_G(g_{\mu\nu}^{(0)}) + S_{EM}(A_\mu^{(0)})$$

- To obtain the correct normalization

$$S_G \rightarrow \frac{1}{G_N} = \frac{\int dx_5}{\hat{G}_N} = \frac{2\pi R}{\hat{G}_N}$$

$$S_{EM} \rightarrow k = \sqrt{16\pi G_N}$$

- Gravity and E.M. unified in higher dimensional space!
- **Geometrical meaning of gauge transf.**
 $d\hat{s}^2$ invariant under local transformations

$$x^5 \rightarrow x^5 - k\Lambda(x), \quad A_\mu^{(0)} \rightarrow A_\mu^{(0)} + \partial_\mu\Lambda(x)$$

Charge Quantization

- Consider scalar field φ coupled to 5-dim gravity

$$S = \int d^5 \hat{x} \sqrt{-\hat{g}} \hat{g}_{MN} (\partial_M \varphi) (\partial_N \varphi)$$

- Each KK mode \mathbf{n} has

$$\text{mass} = \frac{\mathbf{n}}{R} \quad \text{charge} = \frac{\mathbf{n}k}{R}$$

- Determination of fine-structure constant

$$\alpha = \frac{k^2}{4\pi R^2} = \frac{4G_N}{R^2}$$

$$R = \sqrt{\frac{4G_N}{\alpha}} \simeq 4 \times 10^{-31} \text{m} \simeq (5 \times 10^{17} \text{GeV})^{-1}$$

NOT a Theory of real world!

- Φ dynamical fields $\rightarrow \mathbf{F}_{\mu\nu}^{(0)} \mathbf{F}^{(0)\mu\nu} = 0$

In addition to E.M. forces we observe in nature **strong interactions** and **weak interactions**.

- **Weak interactions**

E.g. **beta decay**

$$n \rightarrow p e^- \bar{\nu}$$

is described by Fermi theory of local four-fermion interaction

$$S = \sqrt{2}G_F \int d^4x \left(\bar{\psi}_P \gamma_L^\mu \psi_N \right) \left(\bar{\psi}_\nu \gamma_L^\mu \psi_e \right)$$

where G_F has the same dimension of G_N (in relativistic unities).

$$\sqrt{2}G_F = \frac{1}{\Lambda_{EW}^2} \quad \Lambda_{EW} \simeq 240 \text{ GeV}$$

- We know now that for energies above Λ_{EW} , Fermi theory is described by dynamical **mas-**
sive vectorial gauge fields:

charged (W) and neutral (Z) WEAK gauge bosons. For $|q^2| \ll M_W^2$, $\frac{-i}{q^2 - M_W^2} \rightarrow \frac{i}{M_W^2}$

$$G_F = \frac{g^2}{8M_W^2}$$

Standard Model

- **Unification of E.M. and WEAK interactions** is achieved by **non-abelian gauge theory based on gauge group**

$$SU(2)_L \times U(1)$$

$SU(2)_L$ is spontaneously broken by Higgs mechanism to give masses to W , Z .

- **SM** predictions are in impressive agreement with all precision tests at high energy colliders, e.g. LEP, Tevatron, HERA.

- Unification of gauge couplings at

$$E \simeq 10^{16} \text{ GeV}$$

(achieved in minimal supersymmetric version of SM)

- **Planck mass**: in analogy with Λ_{EW}

$$G_N \simeq \frac{1}{M_{pl}^2} \quad M_{pl} \simeq 10^{19} \text{ GeV}$$

- **Why $\Lambda_{EW} \ll M_{pl}$?**

The Hierarchy Problem

- Two fundamental scales in nature
 - Planck Mass $M_{\text{pl}} \simeq 10^{19} \text{ GeV}$
 - Electroweak scale $\Lambda_{\text{EW}} \simeq 240 \text{ GeV}$
- 17 order of magnitudes of difference !
- Difficult to maintain at quantum level
- Quadratic divergencies in scalar sector

$$(\delta m_H^2 = \alpha \Lambda_{UV}^2 + \dots \text{ fine tuning problem})$$

- Supersymmetry solves this problem

boson \leftrightarrow fermion symmetry

- **Standard approach:** Fundamental theory is at M_{pl} and Λ_{EW} is derived as a small number from dynamics

Q-Gravity in Large Extradimensions

A new approach to the hierarchy problem
(Arkani-Hamed, Dimopoulos, Dvali '98)

Λ_{EW} is a fundamental scale
 M_{pl} is a derived effect from
geometry

This requires

- Compact Extra Spatial Dimensions
- Confinement of matter on subspaces

The idea: Matter fields confined in the usual 4 dimensional space, **Gravity propagating in $D = 4 + \delta$ dimensional space**

Weakness of gravity: Due to large compact spatial dimensions

$$G_N = \frac{G_N^D}{V_{D-4}}$$

Newton law modified at small distances

- $r > R$
$$\mathbf{F}_N(\mathbf{r}) = G_N \frac{m_1 m_2}{r^2}$$

- $r < R$
$$\mathbf{F}_N(\mathbf{r}) = G_N R^\delta \left(\frac{m_1 m_2}{r^{2+\delta}} \right)$$

The potential

$$V(\mathbf{r}) = G_N \frac{m_1 m_2}{r} \left(1 + C e^{-\frac{r}{\lambda}} \right)$$

Plank masses

$$G_N = \frac{1}{M_{\text{pl}}^2}$$

$$G_N^D = \frac{1}{M_D^{D-2}}$$

$$M_{\text{pl}} = \frac{1}{\sqrt{8\pi}} M_D (R M_D)^{\frac{D-4}{2}}$$

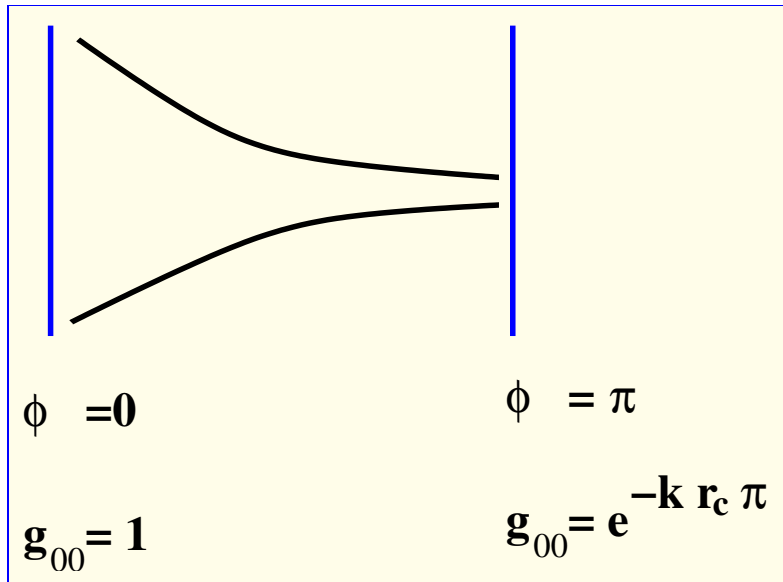
If we want $M_D \simeq 1 \text{ TeV}$, then R must be large in units of M_D^{-1} .

$$D = 4 + \delta, M_D \simeq 1 \text{ TeV} \Rightarrow R \simeq 10^{\frac{30}{\delta} - 17} \text{ cm}$$

- $\delta = 1$ $R = (5 \times 10^{-4} \text{ eV})^{-1} \simeq 10^8 \text{ km}$
- $\delta = 2$ $R = (5 \times 10^{-4} \text{ eV})^{-1} \simeq 0.4 \text{ mm}$

- **Since Newton law has NOT been tested down the mm scale $\Rightarrow \delta \geq 2$**
- **Strong Gravity at TeV scale for $\delta \geq 2$**

Randall-Sundrum Model



- The Einstein Eqs. in 5 dimension, admits a solution for a **non-factorizable metric**

$$ds^2 = e^{-2kr_c|\varphi|} \eta_{\mu\nu} dx^\mu dx^\nu + r_c^2 d\varphi^2$$

where $-\pi < \varphi < \pi$

- Masses related by**

$$\frac{m_\pi}{m_0} = e^{-k r_c \pi}$$

- If compactification radius of 5th dim. is $r_c \simeq 10 \text{ k}^{-1}$

$$\frac{m_\pi}{m_0} \simeq \frac{M_Z}{M_{\text{GUT}}}$$

where M_{GUT} is the scale of Grand Unification $M_{\text{GUT}} \simeq 10^{16} \text{ GeV}$

- Masses subject to a Weyl rescaling of a factor $e^{2kr_c\pi}$.

- If Planck mass in 5-dimensions is of order $M_5 \simeq k \simeq \text{TeV}$

$$M_{\text{pl}}^2 = \frac{M_5^3}{k} [e^{2kr_c\pi} - 1] \gg M_5^2$$

- Hierarchy problem naturally explained for $r_c \cdot k \simeq \mathcal{O}(10)$
- Graviton KK modes have large mass gap and are **strongly** coupled.

Kaluza-Klein excitations of Gravity

- Massive KK Gravitons

$$L_{\text{int}}(\mathbf{x}) \sim \frac{1}{M_{\text{pl}}} \sum_{\mathbf{n}} \mathbf{G}_{\mu\nu}^{(\mathbf{n})}(\mathbf{x}) \mathbf{T}^{\mu\nu}(\mathbf{x})$$

- Massive KK Graviscalars

$$L_{\text{int}}(\mathbf{x}) \sim \frac{1}{M_{\text{pl}}} \sum_{\mathbf{n}} \mathbf{S}^{(\mathbf{n})}(\mathbf{x}) \mathbf{T}_{\mu}^{\mu}(\mathbf{x})$$

$\mathbf{T}_{\mu\nu}(\mathbf{x})$ Energy Momentum Tensor

Masses of KK excitations $M_{\mathbf{n}}^2 \sim \frac{\mathbf{n}^2}{R^2}$

- Impossible to detect **standard quantum gravity** effects for energies of order of $\mathbf{E} < \mathcal{O}(1) \text{ TeV}$
- The production of one single KK mode including zero mode is suppressed by terms of order $(\mathbf{E}/M_{\text{pl}})^2$.
- **Inclusive production of KK modes is huge !**

Inclusive production of KK Gravitons

- Mass splitting

$$\Delta m \simeq \frac{1}{R} = M_D \left(\frac{M_D}{M_{\text{pl}}} \right)^{2/\delta}$$

$$\Delta m = 20\text{KeV} \quad (\delta = 4)$$

$$\Delta m = 7\text{MeV} \quad (\delta = 6)$$

- For small $\delta \leq 6$:

Large number of accessible KK modes.

continuum: $\sum_n \longrightarrow \int dN$

$$dN = A_{\delta-1} \left(\frac{M_{\text{Pl}}^2}{M_D^{2+\delta}} \right) m_G^{\delta-1} dm_G$$

$A_{\delta-1}$ area of δ -sphere, m_G graviton mass

- Inclusive production of KK Gravitons

$$\left(E/M_{\text{pl}} \right)^2 \Rightarrow \left(E/M_D \right)^{2+\delta}$$

Bounds from Astrophys./Cosmology

SuperNova SN1987A collapse

- Nucleon-Nucleon Brehmstrahlung ($T \simeq 30 \text{ MeV}$)

$$\mathcal{L}_{grav} \simeq M_{SN} \times \frac{n_N^3}{\rho} \times (30 \text{ millibarn}) \times \left(\frac{T}{M_D} \right)^{\delta+2}$$

$$\mathcal{L}_{grav} \leq 10^{53} \text{ ergs}^{-1} \simeq (10^{16} \text{ GeV})^2 \Rightarrow M_D > 10^{\frac{15-4.5\delta}{\delta+2}} \text{ TeV}$$

- $M_D > 30 \text{ TeV} (\delta = 2)$ $M_D > 2 \text{ TeV} (\delta = 3)$

Cosmic γ -Rays

$$G \rightarrow \gamma\gamma \quad \tau \simeq 3 \times 10^9 \text{ yrs} \left(\frac{100 \text{ MeV}}{m_G} \right)^3$$

Gravitons sources

- Thermal production after nucleosynthesis :

$$M_D > 110 \text{ TeV} (\delta = 2)$$

$$M_D > 5 \text{ TeV} (\delta = 3)$$

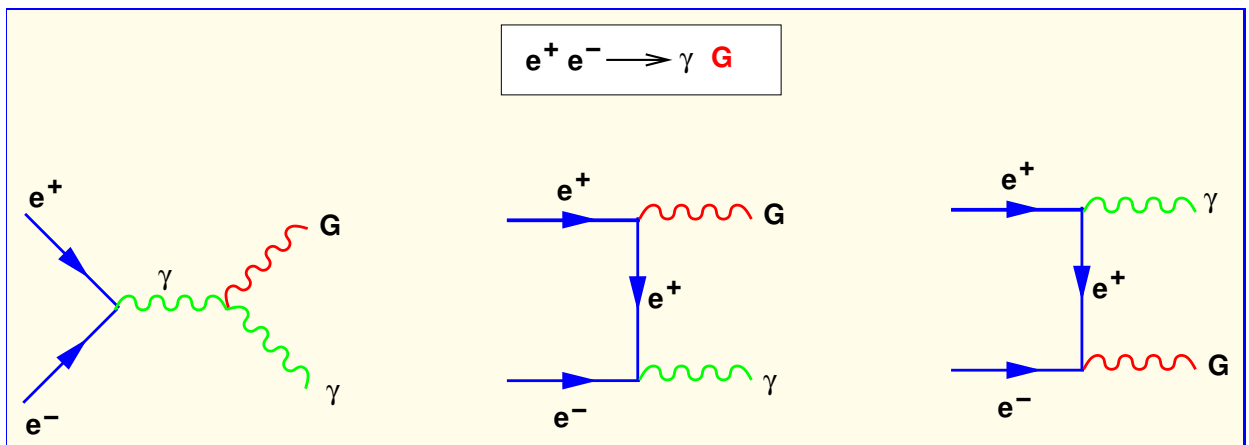
- Neutron stars :

$$M_D > 450 \text{ TeV} (\delta = 2)$$

$$M_D > 30 \text{ TeV} (\delta = 3)$$

Bounds from accelerator experiments

- KK Gravitons production mechanism at e^+e^- colliders



- Massive gravitons G decay outside the detector
- G detected as missing energy
- Hadron colliders (elementary processes)

$$q\bar{q} \rightarrow g + G$$

$$q g \rightarrow g + G$$

$$g g \rightarrow g + G$$

- **Precision Observables:**

- $(g - 2)$ of muon M.L. Graesser (2000)

- precision EW observables at LEP
ADD ('98)

- four-fermion interactions
ADD ('98), G. Giudice, R. Rattazzi, J.D. Wells ('98)

- **Rare decays to Kaluza-Klein Gravitons**

- $\mu \rightarrow \nu_\mu e \bar{\nu}_e + G$ $K \rightarrow \pi + G$

- $J/\Psi \rightarrow \gamma + G$ $Y \rightarrow \gamma + G$

J.Bijnens and M.Maul (2000)

- $Z \rightarrow f\bar{f} + G$ $t \rightarrow bW + G$ $H \rightarrow f\bar{f} + G$

- $H \rightarrow W^+W^- + G$

E.G. and B.Mele (2002)

Experimental Bounds on M_D

G. Giudice, R. Rattazzi, and J.D. Wells ('98)

discovery requirement ($\sigma_{\text{Signal}} > 5\sqrt{\sigma_B \mathcal{L}}/\mathcal{L}$)

$$e^+e^- \rightarrow \gamma + G$$

G as missing energy

- **LEP200**, $\mathcal{L} = 4 \times 500 \text{ fb}^{-1}$ ($\sqrt{s}=200 \text{ GeV}$)

$$\delta = 2 \quad M_D > 1.3 \text{ TeV}$$

- **e^+e^- LC** $\mathcal{L}=200 \text{ fb}^{-1}$ ($\sqrt{s}=1 \text{ TeV}$)

$$\delta = 2 \quad M_D > 4.1 \text{ TeV}$$

$$pp \rightarrow \text{Jets} + G$$

G as missing energy

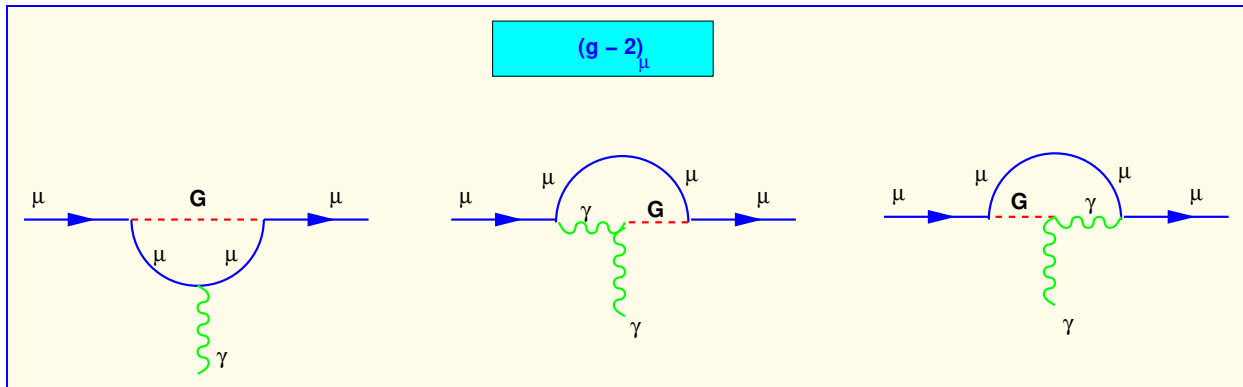
- **LHC**, $\mathcal{L} = 100 \text{ fb}^{-1}$

$$\delta = 2 \quad M_D > 8.5 \text{ TeV}$$

Bounds from precision observables

$g-2$ of the muon

Graesser (2000)



- It is very sensitive to new physics
- Contribution of each KK graviton to $g-2$ is finite. \sum over KK modes is divergent!

$$\Delta a_\mu = \frac{7}{12\pi} \left(\frac{m_\mu}{M_D} \right)^2 \times \lambda \quad (\delta = 2)$$

$$\Delta a_\mu^{\text{Ext. Dim}} = 2 \times 10^{-9} \left(\frac{\text{TeV}}{M_D} \right)^2 \times \lambda$$

BNL E821 $\Rightarrow \Delta a_\mu < 4 \times 10^{-10}$ at 2σ level

$$M_D > 2.2 \text{ TeV} \sqrt{\lambda}$$

Rare decays to KK gravitons

- Decay $\boxed{\mathbf{K} \rightarrow \pi + \mathbf{G}}$ ADD ('98)

Due to angular momentum conservation \Rightarrow spin-0
KK gravitons

$$\Gamma_{K \rightarrow \pi + G} \simeq \frac{1}{16\pi} \left(\frac{m_K^5}{m_W^4} \right) \times \left(\frac{m_K}{M_D} \right)^{\delta+2}$$

$$\boxed{\text{BR} < 10^{-12}} \quad (\text{for } \delta = 2 \text{ and } M_D = 1 \text{ TeV})$$

- $\boxed{\text{Quarkonium}(q\bar{q}) \rightarrow \gamma + \mathbf{G}}$

Bijnens and Maul (2000)

Y decays are more favorable than $J/\psi \rightarrow \gamma + G$

$$\boxed{Y \rightarrow \gamma + \mathbf{G}} \text{ spin-2 dominated}$$

$$\text{BR}^{\text{exp}}(Y \rightarrow \gamma + G) < 10^{-5}$$

$$\boxed{M_D > 50 \text{ (9) GeV}} \text{ for } \delta = 2 \text{ (6)}$$

Bounds on M_D from Z, W, t, H decays

E.G. and B.Mele, Nucl.Phys. B647, 319 (2002)

$$\text{BR}(Z \rightarrow \sum_f f\bar{f} + G_X) = 8.2 \times 10^{-8} \left(\frac{\text{TeV}}{M_D} \right)^4 \quad (\delta = 2)$$

$$\text{BR}(W \rightarrow \sum_{f',f} f'\bar{f} + G_X) = 5 \times 10^{-8} \left(\frac{\text{TeV}}{M_D} \right)^4 \quad (\delta = 2)$$

$$\text{BR}(t \rightarrow W b + G_X) = 1.8 \times 10^{-7} \left(\frac{\text{TeV}}{M_D} \right)^4 \quad (\delta = 2)$$

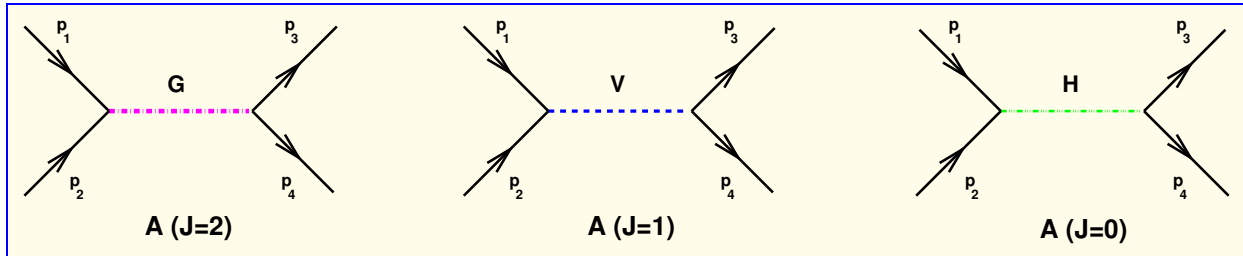
Higgs decays $(\delta = 2)$ m_H

$$\text{BR}(H \rightarrow \bar{b}b + G_X) = 2.2 \times 10^{-7} \left(\frac{\text{TeV}}{M_D} \right)^4 \quad 120 \text{ GeV}$$

$$\text{BR}(H \rightarrow WW + G_X) = 2.1 \times 10^{-5} \left(\frac{\text{TeV}}{M_D} \right)^4 \quad 500 \text{ GeV}$$

Virtual KK Gravitons exchange

Virtual KK Gravitons exchange affect the SM processes

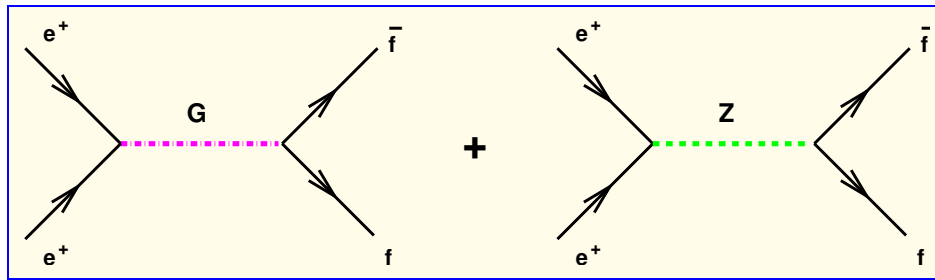


$e^-e^+ \rightarrow f\bar{f}$, $e^-e^+ \rightarrow \gamma\gamma$, $W^+W^- \rightarrow W^+W^-$, etc.

- **Real part of amplitude $\text{Re}[A]$ is divergent.**
Predictivity in terms of M_D is lost.
- **Imaginary part $\text{Im}[A]$ is finite** (due to branch-cut singularity of real gravitons emission)
- **Resonant SM processes** can have interference with **$\text{Im}[A]$** of the graviton exchange
Finite result predicted in terms of M_D .
- **Virtual Graviton** exchanges at the Z-pole in large Extra Dimensions
A. Datta, E.G, and B. Mele, PLB 552 (2003) 237

Gravitational interference at Z pole

A. Datta, E.G, and B. Mele, PLB 552 (2003) 237



$$\frac{d\sigma_f}{d\cos\theta} = \frac{9\pi}{2M_Z^2} \left(\frac{\Gamma_e\Gamma_f}{\Gamma_Z^2} \right) S_f(\cos\theta)$$

$$S_f(x) = 1 + x^2 - \Delta_1 (1 - 3x^2) + 2A_e A_f (x + \Delta_2 x^3)$$

$$\Delta_{1,2} \propto \left(\frac{M_Z}{M_D} \right)^{2+\delta}$$

- **New contribution: vanishes** when integrated over the full range of $\cos\theta = (-1, 1)$
- $\delta = 2$, $M_D = 1 \text{ TeV} \Rightarrow \Delta_{1,2} \simeq \mathcal{O}(10^{-4})$
too small for present **EXP.** sensitivity

Virtual KK Gravitons in Higgs production

A. Datta, E.G, and B.Mele, JHEP 0310:003,2003

Light Higgs

- **Hadron Colliders:** $gg \rightarrow H \rightarrow \gamma\gamma$ (1-loop)²
- Interference with $Im[A(gg \rightarrow G, S \rightarrow \gamma\gamma)]$
- Quite small effect, being coupled to the $Trace[T_\mu^\mu]$

For massless final states $T_\mu^\mu \propto \mathcal{O}(\alpha_S, \alpha_e)$
gravitational anomalies.

Heavy Higgs

- **e^+e^- Colliders:** $WW \rightarrow H \rightarrow WW$ Tree!
- Large interference $Im[A(WW \rightarrow G, S \rightarrow WW)]$
since $T_\mu^\mu \propto M_W^2$
- **$\mu^+\mu^-$ Colliders:** $\mu^+\mu^- \rightarrow H \rightarrow (WW), (\bar{t}t)$

Conclusions

- Scenarios of large Extra Dimensions offer alternative ways to solve the Hierarchy problem
- If QG scale (M_D) is in the TeV region, dramatic scenarios with non standard physics are expected at LHC:
strong gravity at TeV scale, black holes formations.
- Negative search at LEP200 $e^+e^- \rightarrow \gamma + G$ set bounds of order $M_D \geq 1$ TeV for 2 extra dimensions, comparable with $Z \rightarrow f\bar{f} + G$ from LEP1.
- In perturbative region, **LHC can probe QG scale up to 8 TeV for $\delta = 2$**